

The Commander parameter file

Beyond PLANCK

Kristian Joten Andersen

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What is the Commander parameter file?

- All information going into any Commander run
 - The "recipe"

- Three main parameter types
 - Infrastructure parameters; these control the behaviour of the algorithms and IO
 - Data set parameters; properties of each data set
 - Model parameters; properties of each sky component
- Commander documentation is available through the project home page: <u>https://docs.beyondplanck.science/#/README</u> (direct link)
- Caveats:
 - This is a software platform for cutting-edge research, and therefore by nature a continuous work-in-progress.
 - "Alpha state"
 - Support is provided on a strictly voluntary basis; there is no designated "help desk"
 - If you find information missing, please contact us.



General notes about read-in

- Any line starting with # is assumed as a comment
- Blank lines are allowed

- Any specific path should be given inside apostrophes
 - 'path/to/some/data/data.dat'
- Every parameter takes the form:
 PARAMETER_NAME = value
- The *first* occurance of a parameter is the one that will be read

General notes about read-in

@INCLUDE

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- Recurse into the given file
- "@INCLUDE parfile2.txt" will include parameters from parfile2.txt at the current position in the original parameter file

Command line read-in

- Any parameter may be specified using the command line
- --PARAMETER_NAME=value (no spaces)
- Will override the parameter file for the specified parameter



Error checking and debugging

Parameter file is parsed at the beginning of the run

- Stored in-memory in a Fortran type called "cpar" ("Commander parameters")
- Active filenames are validated with respect to existence, but not content.
- Many types of parameter file errors are therefore (automatically) detected immediately, but not all.



Error checking and debugging

If you find that the code crashes with something that looks like a parameter error, typical things to check are the following:

• Is a given parameter of the correct/expected type?

- Check which line causes the crash, and look it up in comm_param_mod.f90
- Does a given file contain the expected data type?
- If it is an ASCII input file, does it have the correct format?





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Algorithm specification

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= parameters that one is likely to change on a (semi-)regular basis when working with Commander, see the documentation for details.

	OPERATION VERBOSITY	= sample = 3	# {sample,optimize} # [0,,3]	
	######################################	############ specifica ####################################	######################################	
$\begin{array}{c} \rightarrow \\ \rightarrow \\ \rightarrow \\ \rightarrow \\ \rightarrow \end{array}$	<pre># Monte Carlo options NUMCHAIN NUM_GIBBS_ITER BASE_SEED CHAIN_STATUS NUM_INIT_CHAINS INIT_CHAIN01 NUM_GIBBS_STEPS_PER_TOD_SAMPL</pre>	= 1 = 1000 = 163425 = append = 1 = "data_B path E = 1 How	<pre># Number of independent chains # Length of each Markov chain # Seed for random number generat {append, new} P8/chain_init_BP8.15.h5:9" {new name sample y often to run TOD analysis</pre>	or }
* * * *	SAMPLE_ONLY_POLARIZATION SAMPLE_SIGNAL_AMPLITUDES SAMPLE_SPECTRAL_INDICES SAMPLE_POWSPEC	= .false. = .true. = .true. = .false.	Component separation analysis	
→ • *	ENABLE_TOD_ANALYSIS TOD_OUTPUT_4D_MAP_EVERY_NTH_I TOD_OUTPUT_AUXILIARY_MAPS_EVE TOD_INCLUDE_ZODI FFTW3_MAGIC_NUMBERS TOD_NUM_BP_PROPOSALS_PER_ITER	= .true. TER = 10 RY_NTH_ITE = .false. = 'data_B = 1 # 1	R = 10 P8/fft3_magic_numbers_230810.txt' L for sampling; >= 1 for optimize	TOD
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Algorithm specification

<pre># Options for CMB resampling (RESAMPLE_CMB FIRST_SAMPLE_FOR_CMB_RESAMP LAST_SAMPLE_FOR_CMB_RESAMP NUM_SUBSAMP_PER_MAIN_SAMPLE</pre>	for = 1 = 1 = 1	constraine false. .5 .0	∍d	realization production)
# Numerical accuracy settings				
G CONVERGENCE CRITERION	=	fixed iter	#	{residual. chisquare}
CG LMAX PRECOND	=	-1	#	lmax for low-l preconditioner
CG MAXITER	=	300	#	Conjugate gradients time out limit
CG ⁻ MINITER	=	5		
CG_TOLERANCE	=	1.d-8	#	Fractional CG convergence criterion
CG_CONV_CHECK_FREQUENCY	=	1	#	Check convergence every n'th iteration
CG_PRECOND_TYPE	=	diagonal	#	{diagonal, pseudoinv} Seljebotn et al. 2019
CG_INIT_AMPS_ON_ZERO	=	.false.		
<pre>SET_ALL_NOISE_MAPS_TO_MEAN</pre>	=	.false.		
NUM_INDEX_CYCLES_PER_ITERATION	=	1		
IGNORE GAIN AND BANDPASS CORR	= .	false.		

* If using chisq: set to ~5 as this check is more expensive than residual

** Set all component amplitudes to zero for each CG search. Can lead to long convergence time



Output options

→ OUTPUT DIRECTORY = chains BP8 c16 Path to where all output files are written THINNING FACTOR = NSIDE CHISQ = 16POLARIZATION CHISQ = .true. OUTPUT MIXING MATRIX = .false. OUTPUT RESIDUAL MAPS = .true. OUTPUT CHISQ MAP = .true. OUTPUT_EVERY_NTH_CG_ITERATION = 0 OUTPUT_CG_PRECOND_ETGENVALS = .false. OUTPUT INPUT MODEL = .false. OUTPUT DEBUG SEDS = .false. OUTPUT SIGNALS PER BAND = .false.

Different options of what to output

- Output each instance as a HEALPix map at the band resolution/pixelarization. Use with care as this will require a lot of disk space.
- ** Outputs non-converged CG iterations to files. Use with care as they will overwrite converged samples with the same numbering. Only use for debugging.



Data sets: inclusion

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	# Da	ata se	ets		#
	#######################################	#####	*#########	#############	##########
			data DDO		Path to where (almost) all input data is stored
\rightarrow	NUMBAND	=	14		How many frequency bands are defined (any band with # higher will not be read)
	# LFI				
→	INCLUDE_BAND001	=	.true.	# 30 GHz	Per band inclusion flags.
	INCLUDE_BAND002	=	.true.	# 44 GHz	
	INCLUDE_BAND003	=	.true.	# 70 GHz	
	# HFI T				
	INCLUDE_BAND004		.true.	# 857 GHz	
	# Haslam and WMAP T				
	INCLUDE_BAND005		.true.	# Haslam	
	INCLUDE_BAND006	=	.true.	# WMAP Ka	Т
	INCLUDE_BAND007	=	.true.	# WMAP Q1	Т
	INCLUDE_BAND008	=	.true.	# WMAP Q2	Т
	INCLUDE_BAND009		.true.	# WMAP V1	Т
	INCLUDE_BAND010		.true.	# WMAP V2	Т
	# HFI P				
	INCLUDE_BAND011	=	.true.	# 353 GHz P	
	# WMAP P				
	INCLUDE_BAND012	=	.true.	# WMAP Ka P	
	INCLUDE_BAND013	=	.true.	# WMAP Q P	
	INCLUDE_BAND014		.true.	# WMAP V P	



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Data sets: smoothing scales

	SOURCE_MASKFILE PROCESSING_MASKFILE PROCESSING_MASKFILE2 PROC_SMOOTH_SCALE	= none = none = none = 30.	<pre>#bright_sources.txt #procmask.fits #procmask2.fits #arcmin; smoothing inside processing mask</pre>
	<pre># Spectral index sampling optic NUM_SMOOTHING_SCALES</pre>	ons = 1	
**	SMOOTHING_SCALE_FWHM01 SMOOTHING_SCALE_LMAX01 SMOOTHING_SCALE_NSIDE01 SMOOTHING_SCALE_PIXWIN01	= 300. = 96 = 32 = pixel	# _window_n0032.fits HEALPix pixel window file
***	SMOOTHING_SCALE_FWHM_POSTPROC01	L = 600.	# Smoothing FWHM after sampling

- Helpful to mask out problematic point sources / regions, read documentation for details.
- ****** [arcmin] larger than any FWHM of bands included in sampling at this scale

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*** [arcmin] Any sampled spectral parameter at this smoothing scale will be smoothed with a beam of this FWHM after sampling. For pixel-region sampling, this is the smoothing between regions!



Data sets: band specifics

General parameter syntax: BAND_PARAMETERxxx = value xxx = 3-digit band number

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How to add new data bands will be shown in tomorrow's tutorials

	# 30 GHz parameters		
→	BAND_LABEL001	=	030
	BAND_TOD_TYPE001	=	'LFI' #{LFI, WMAP, none}
	BAND_OBS_PERIOD001	=	1
	BAND_POLARIZATION001	=	.true.
	BAND_NSIDE001	=	512
	BAND_LMAX001	=	1500
	BAND_UNIT001	=	uK_cmb
	BAND_NOISE_FORMAT001	=	rms
→	BAND_MAPFILE001	=	BP_030_map.fits
⇒	BAND_NOISEFILE001	=	BP_030_rms.fits
→	BAND_REG_NOISEFILE001	=	none # BP_030_regnoise.fits
⇒	BAND_NOISE_RMS001_SMOOTH01	=	BP_030_rms_smoothscale1.fits {native, none}
	BAND_NOISE_UNIFORMIZE_FSKY001	=	0.0
	BAND_MASKFILE001	=	fullsky
	BAND_BEAMTYPE001	=	b_l # {b_l, febecop}
	BAND_BEAM_B_L_FILE001	=	Bl_TEB_npipe6v19_30GHzx30GHz.fits
	BAND_BEAM_B_PTSRC_FILE001	=	febecop_AT20G_GB6_NVSS_PCCS2_v6_030.h5
	BAND_PIXEL_WINDOW001	=	pixel_window_n0512.fits
	BAND_SAMP_NOISE_AMP001	=	.false.



Data sets: band specifics

	BAND_SAMP_BANDPASS001	=	.false.
	BAND_BANDPASSFILE001	=	LFI_instrument_v4.h5
→	BAND_SAMP_GAIN001	=	.false.
⇒	BAND GAIN PRIOR MEAN001	=	= 1.
⇒	BAND GAIN PRIOR RMS001	Ē	= 0.1
→	BAND_GAIN_CALIB_COMP001	=	all 'cmb'
	BAND_GAIN_LMIN001	=	-1
	BAND_GAIN_LMAX001	=	-1 < 0 ~ no uppernower innit
	BAND_GAIN_APOD_MASK001	=	fullsky
	BAND_GAIN_APOD_FWHM001	=	120.
→	BAND_MASKFILE_CALIB001	=	<pre>mask_common_dx12_n0512_TQU.fits</pre>
	BAND_DEFAULT_GAIN001	=	1
	BAND_DEFAULT_BP_DELTA001	=	0.
	BAND_DEFAULT_NOISEAMP001	=	1.
	BAND COMPONENT SENSITIVITY001	=	broadband



Data sets: band specifics

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	BAND_TOD_MAIN_PROCMASK001 BAND_TOD_SMALL_PROCMASK001	<pre>= mask_proc_030_res_v = mask_smap6.fits</pre>	v5.fits
	BAND TOD BP INIT PROP001	= bp init 030 v1.dat	
	BAND TOD RIMO001	= LFI instrument v4.1	h5
*	BAND_TOD_FILELIST001	= filelist_30_v17.tx	t
	BAND_TOD_HALFRING001	= 0	
	BAND_TOD_START_SCANID001	= 3	
	BAND_TOD_END_SCANID001	= 44072	
	BAND_TOD_TOT_NUMSCAN001	= 45860	
	BAND_TOD_FLAG001	= 6111232	
	BAND_TOD_ORBITAL_ONLY_ABSCAL001	l =.false.	
	BAND_TOD_DETECTOR_LIST001	= "27M,27S,28M,28S"	
→	BAND_TOD_INIT_FROM_HDF001	= default	{default, none, path to chain+sample}

* TOD scan definition file. List of all compressed time-ordered data files with initial values. See documentation.



****	*	G	obal model pa	rameter	s	
European Commissi	on	MJYSR_CONVENTION T_CMB	= IRAS = 2.7255d0			
	→ →	INSTRUMENT_PARAM_FILE INIT_INSTRUMENT_FROM_HDF	= instrument_params.da = default	at	Initial band gain a bandpass correc	and tions
		CMB_DIPOLE_PRIOR # 'mask_common_dx12_n1024_TQU	= none J.fits; 3364.4; 263.998;	; 48.265' #	LFI 2018	
	* *	NUM_SIGNAL_COMPONENTS INCLUDE_COMP01 INCLUDE_COMP02 INCLUDE_COMP03 INCLUDE_COMP04	= 4 = .true. # Cmb # CMB; = .true. # synch # Sy = .true. # dust # The = .true. # md # Mono	; no monopol /nch pow-law ermal dust and dipoles	e {.true., .false.}	
		OUTPUT_COMPS_TO_CHAINDIR = 'a	11'			
CG	† †††	<pre>NUM_CG_SAMPLING_GROUPS = 2 CG_SAMPLING_GROUP01 CG_SAMPLING_GROUP_MASK01 CG_SAMPLING_GROUP_MAXITER01 CG_SAMPLING_GROUP02 CG_SAMPLING_GROUP_MASK02 CG_SAMPLING_GROUP_MAXITER02</pre>	<pre>= 'md' = mask_common_dx12_n102 = 3 = 'cmb,dust,synch' = fullsky = 50</pre>	24_TQU.fits Component Sampling m Max CG iter	labels ask rations (# of iterations	s for fixed_iter)
		<pre># Alm sampler settings ALMSAMP_NSAMP_ALM = 1 ALMSAMP_BURN_IN = 1 ALMSAMP_NSIDE_CHISQ_LOWRES ALMSAMP_PRIOR_FWHM ALMSAMP_OPTIMIZE_ALM ALMSAMP_OPTIMIZE_ALM ALMSAMP_PPLY_PRIOR ALMSAMP_PIXREG ALMSAMP_PRIORSAMP_FROZEN_REGI</pre>	00 # of mcmc samples p # of gibbs iterati = 16 = 0 = .false. # s = .true. # ap = .true. SONS = .true.	per gibbs ions with st save chisq f oply prior t	eplength adjustment rom prev gibbs iter o alms	a_Im spectral parameter sampler
		<pre>#local sampler settings LOCALSAMP_BURN_IN = 1</pre>	# of gibbs iterations 16	with steple	ngth adjustment	Beyond PLANCK

Model specific parameters

European Commission	# CMD			
→	# CMB COMP_LABEL01 COMP_TYPE01 COMP_CLASS01 COMP_POLARIZATION01 COMP_CG_SCALE01	<pre>= cmb = cmb = diffuse # {diffuse, ptsrc, te = .true. = 1.d0</pre>	General para COMP_PAR/ mplate} xx = 2-digit b	meter syntax: AMETERxx = value and number
	COMP_NSIDE01 COMP_MONOPOLE_PRIOR01 COMP_DEFLATION_MASK01 COMP_L_APOD01 COMP_LMIN_AMP01 COMP_LMAX_AMP01 COMP_LMAX_IND01	<pre>= 1024 = none "monopole:mask_common_dx1 = fullsky = 2000 = 0 = 2000 = 0</pre>	2_n1024_TQU.fits"	
	COMP_OUTPUT_FWHM01 COMP_UNIT01 COMP_NU_REF_T01 COMP_NU_REF_P01 COMP_CL_TYPE01 COMP_CL_POLTYPE01 COMP_CL_BETA_PRIOR_MEAN01 COMP_CL_BETA_PRIOR_RMS01 COMP_CL_DEFAULT_AMP_T01	<pre>= 14 # arcmin = uK_cmb = 1 100. = 1 100. = power_law # {none, single_l, bi = 1 # {1 = {T+E+B}, 2 = {T,E+B}, = 0.0 = 0.1 = 20 # Pivot multipole = 1000000 # D l = amp * (l</pre>	nned, power_law} 3 = {T,E,B}} /lpivot)**beta	
prior RMS	COMP_CL_DEFAULT_AMP_E01 COMP_CL_DEFAULT_AMP_B01 COMP_CL_DEFAULT_BETA_T01 COMP_CL_DEFAULT_BETA_E01 COMP_CL_DEFAULT_BETA_B01 #COMP_CL_TYPE01 #COMP_CL_BIN_FILE01 #COMP_CL_DEFAULT_FILE01 COMP_MASK01 COMP_INPUT_AMP_MAP01 COMP_PRIOR_AMP_MAP01 COMP_OUTPUT_ER_MAP01	<pre>= 1000 = 1000 = 0.d0 = -0.5d0 = binned # {none, binned, pow = bins_lmax2000_TE.dat # for bi = base_plikHM_TTTEEE_lowl_lowE_le = fullsky = init_cmb_amp_BP8.11.fits initialit = none</pre>	er_law} nned type nsing.minimum.theory_cl zation map prior	
→	COMP_DOTFOT_EB_MAPO1	= default		

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Model specific parameters

sion	<pre># Synchrotron component</pre>		
	COMP_LABEL02	=	synch
	COMP_TYPE02	=	power_law
	COMP_CLASS02	=	diffuse # {diffuse, ptsrc}
	COMP_POLARIZATION02	=	.true.
→	COMP_CG_SCALE02	=	1 Multiplicative amplitude scale used in CG search
→	COMP_CG_SAMP_GROUP_MAXITER02	=	40 Max CG iterations in amplitude sampling after marginal
	COMP_NSIDE02	=	1024 likelihood spectral index sampling
	COMP_MONOPOLE_PRIOR02	=	none
	COMP_DEFLATION_MASK02	=	fullsky
	COMP_L_APOD02	=	1500
	COMP_LMIN_AMP02	=	0
	COMP_LMAX_AMP02	=	1500
	COMP_OUTPUT_FWHM02	=	60 # arcmin
	COMP_UNIT02	=	uK_RJ
→	COMP_NU_REF_T02	=	30
	COMP_NU_REF_P02	=	30
	COMP_MASK02	=	fullsky
	COMP_CL_TYPE02	=	gauss # {none, single_l, binned, power_law}
	COMP_CL_POLTYPE02	=	$2 \# \{1 = \{T+E+B\}, 2 = \{T,E+B\}, 3 = \{T,E,B\}\}$
	COMP_CL_BETA_PRIOR_MEAN02	=	-0.5
	COMP_CL_BETA_PRIOR_RMS02	=	0.1
	COMP_CL_L_PIVOT02	=	100 # Pivot multipole
	COMP_CL_DEFAULT_AMP_T02	=	1e3
	COMP_CL_DEFAULT_AMP_E02	=	200
	COMP_CL_DEFAULT_AMP_B02	=	100
	COMP_CL_DEFAULT_BETA_T02	=	60d0
	COMP_CL_DEFAULT_BETA_E02	=	3000
	COMP_CL_DEFAULT_BETA_B02	=	30d0



Model specific parameters: Synchrotron

→ COMP INDMASK02 COMP LMAX IND02 COMP PRIOR UNI BETA LOW02 COMP PRIOR UNI BETA HIGH02 → COMP PRIOR GAUSS BETA MEAN02 → COMP PRIOR GAUSS BETA RMS02 COMP APPLY JEFFREYS PRIOR02 → COMP BETA SMOOTHING SCALE02 → COMP BETA POLTYPE02 COMP BETA NU MIN02 COMP BETA NU MAX02 COMP INPUT AMP MAP02 COMP PRIOR AMP MAP02 COMP INPUT BETA MAP02 COMP DEFAULT BETA02 COMP OUTPUT EB MAP02

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COMP_INIT_FROM_HDF02

= 100= -4.5 = -1.5 = -3.1 Prior mean = 0.1 Prior RMS = .false. = 2 = 2 # index $\{1 = \{T+Q+U\}, 2 = \{T,Q+U\}, 3 = \{T,Q,U\}\}$ = 0. # Lowest frequency for index estimation in GHz = 80. # Highest frequency for index estimation in GHz = init synch amp.fits = none = init synch beta.fits default = -3.1= .false.

= mask synch beta BP8 10deg new chisqmask.fits

= default



Model specific parameters

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⇒	COMP_BETA_INT_LMAX02	=	-1 #	alm sampling (>=0), local sampling (-1)
	COMP_BETA_INT_LNLTYPE02	=	marginal	<pre># {chisq,ridge,marginal,prior}</pre>
	COMP_BETA_INT_PIXREG02	=	fullsky	# {fullsky,single_pix,pixreg}
	COMP_BETA_INT_SAMPLE_NPROP02	=	.false.	
	COMP_BETA_INT_SAMPLE_PROPLEN02	2 =	.true.	
	COMP_BETA_NPR0P02	=	fullsky	<pre># nprop map, local sampling (fullsky = 1)</pre>
	COMP_BETA_INT_NPROP_INIT02	=	1000	<pre># {> 0, < 0 to disable}. overwrites nprop</pre>
				<pre># init values from nprop map. local sampler</pre>
	COMP_BETA_UNI_NPROP_LOW02	=	10	<pre># {>= 0} local sampling. minimum number</pre>
				# of proposals per pixel region
	COMP_BETA_UNI_NPROP_HIGH02	=	2000	<pre># {> 0} local sampling. minimum number</pre>
				<pre># of proposals per pixel region</pre>
	COMP_BETA_MASK02	Ξ	<pre>mask_syn</pre>	ch_beta_local.fits
	COMP_BETA_PROPLEN02	=	fullsky	<pre># proposal length map, local sampling</pre>
				# (fullsky = 1.d0)
	COMP_BETA_INT_PROPLEN_INIT02	=	3.d-3	<pre># {> 0, < 0 to disable} overwrites proplen</pre>
				# init values from map
	COMP_BETA_ALMSAMP_INIT02	=	init_alm	_synch_beta_9reg.dat
	COMP_BETA_INT_NUM_PIXREG02	=	9 # numb	er of pixel regions to sample (from 1 to N)
			# regi	ons above N set to O (and prior value)
	COMP_BETA_INT_FIX_PIXREG02	=	none	<pre># {none, '1,3,4'} pixel regions to fix,</pre>
				# i.e. freeze on init
	COMP_BETA_INT_PIXREG_PRIORS02	=	none #	<pre>{none, string with prior means of all</pre>
				pixel regions}
	COMP_BETA_PIXREG_MAP02	=	<pre>map_9reg</pre>	ions_n1024.fits
				<pre>#(from 1 -> N). 'fullsky' -> all pixels = 1</pre>
	COMP_BETA_PIXREG_INITVALUE_MAI	P02	$2 = init_{}$	synch_beta_pixreg.fits # {none, mapname}

Many of these parameters are depending on each other. See documentation for details!



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Model specific parameters

** * → COMP BETA INT LMAX02 = -1 # alm sampling (>=0), local sampling (-1) COMP BETA INT LNLTYPE02 = marginal # {chisg,ridge,marginal,prior} COMP BETA INT PIXREG02 # {fullsky,single pix,pixreg} = fullsky COMP BETA INT SAMPLE NPROP02 = .false. COMP BETA INT SAMPLE PROPLEN02 = .true. COMP BETA NPROP02 = fullsky # nprop map, local sampling (fullsky = 1) COMP BETA INT NPROP INIT02 $\# \{ > 0, < 0 \text{ to disable} \}$. overwrites nprop = 1000 # init values from nprop map. local sampler = 10 COMP BETA UNI NPROP LOW02 # {>= 0} local sampling. minimum number # of proposals per pixel region = 2000 # {> 0} local sampling. minimum number COMP BETA UNI NPROP HIGH02 # of proposals per pixel region = mask synch beta local.fits # local sampler mask COMP BETA MASK02 COMP BETA PROPLEN02 = fullsky # proposal length map, local sampling # (fullsky = 1.d0)COMP BETA INT PROPLEN INITO2 = 3.d-3 # {> 0, < 0 to disable} overwrites proplen # init values from map COMP BETA ALMSAMP INIT02 = init alm synch beta 9reg.dat COMP BETA INT NUM PIXREG02 = 9 # number of pixel regions to sample (from 1 to N) # regions above N set to 0 (and prior value) COMP BETA INT FIX PIXREG02 # {none, '1,3,4'} pixel regions to fix, = none # i.e. freeze on init COMP_BETA_INT_PIXREG_PRIORS02 = none # {none, string with prior means of all pixel regions} = map 9regions n1024.fits # Pixel region map COMP BETA PIXREG MAP02 $#(from 1 \rightarrow N)$. 'fullsky' \rightarrow all pixels = 1 COMP BETA PIXREG INITVALUE MAP02 = init synch beta pixreg.fits # {none, mapname} * Affects input spectral parameter map. If pixel-by-pixel structure, this must be -1 ****** Defined for poltype (polarization type), "INT" = poltype index 1 {T or T+Q+U},

"POL" = poltype index 2 {Q+U or Q}, "POL3" = poltype index 3 {U}



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"BeyondPlanck"

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Questions?

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Cosmoglobe Beyond